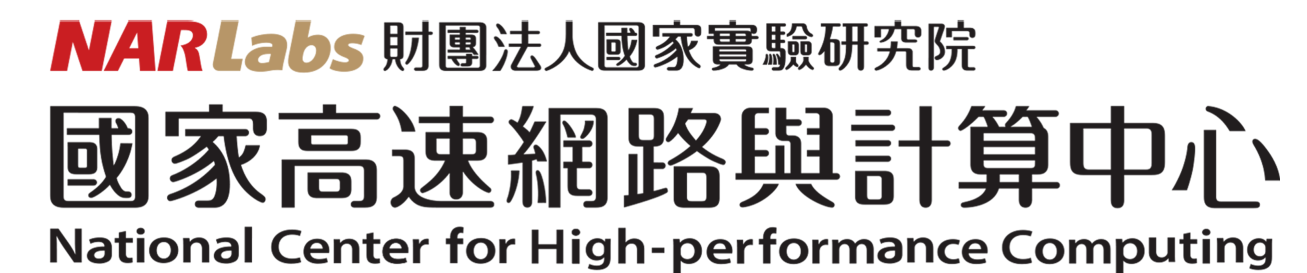


# Streaming instabilities in modern protoplanetary disks

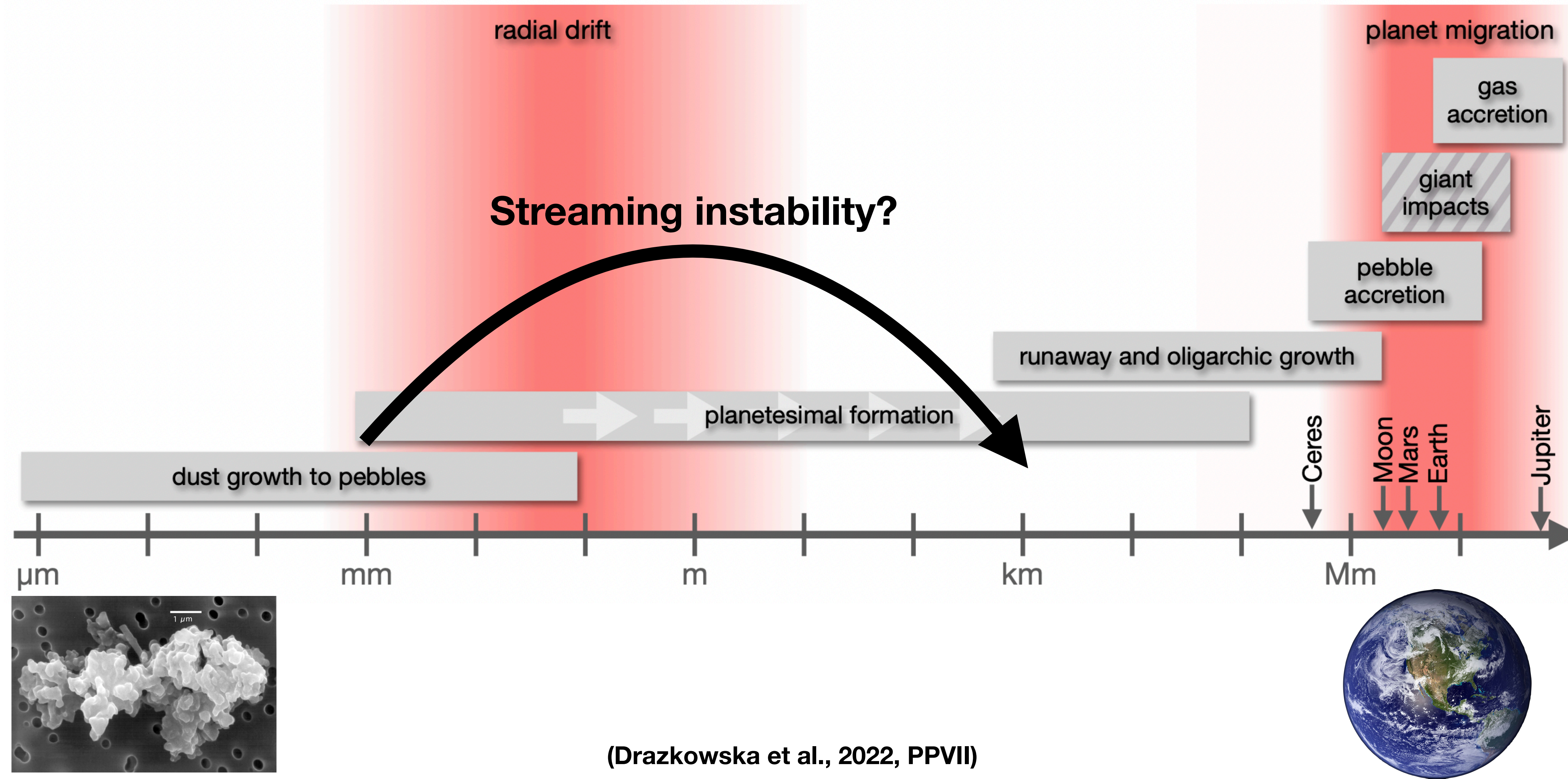
Min-Kai Lin

September 2022



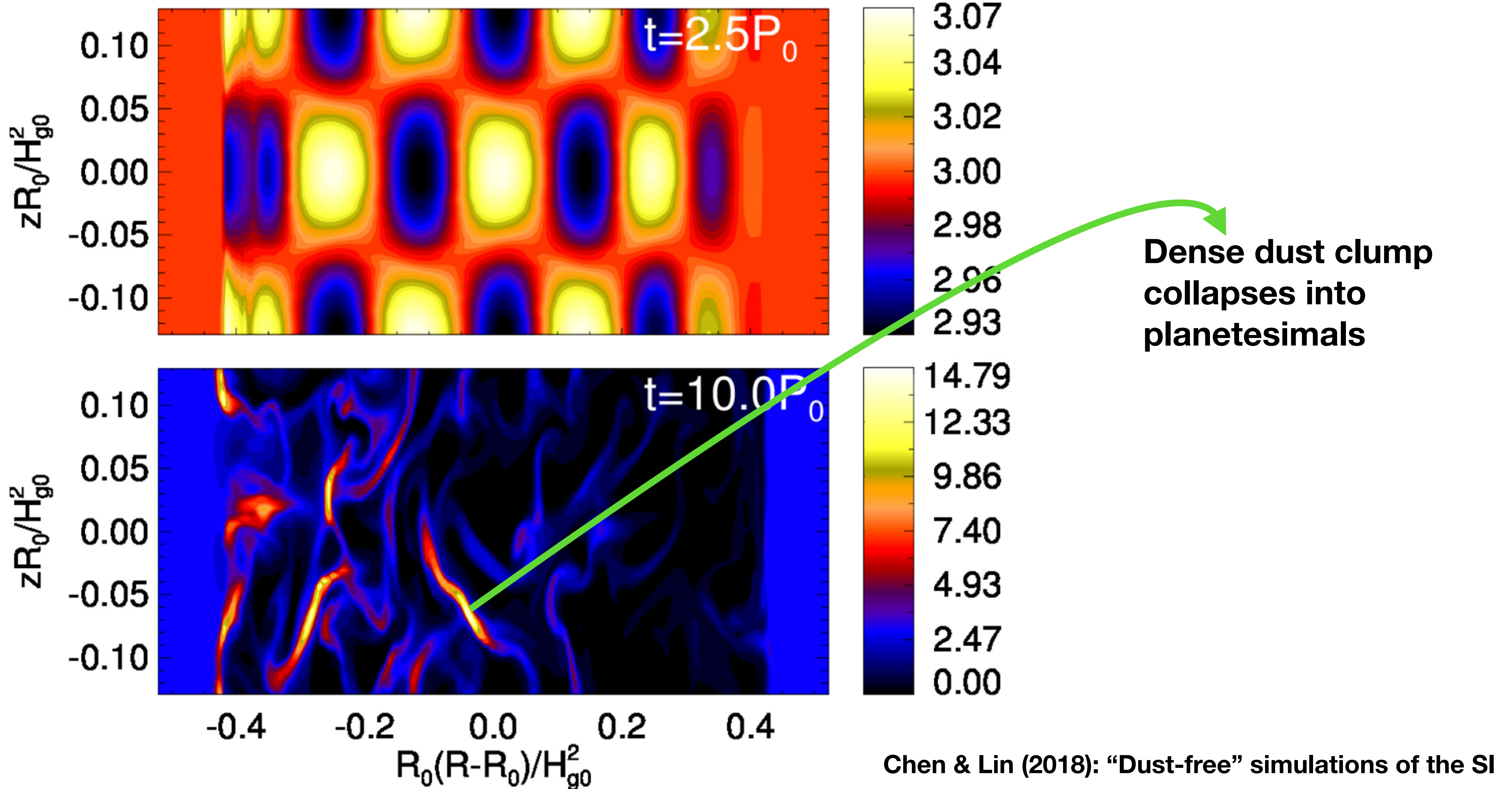


# Conventional planet formation theory

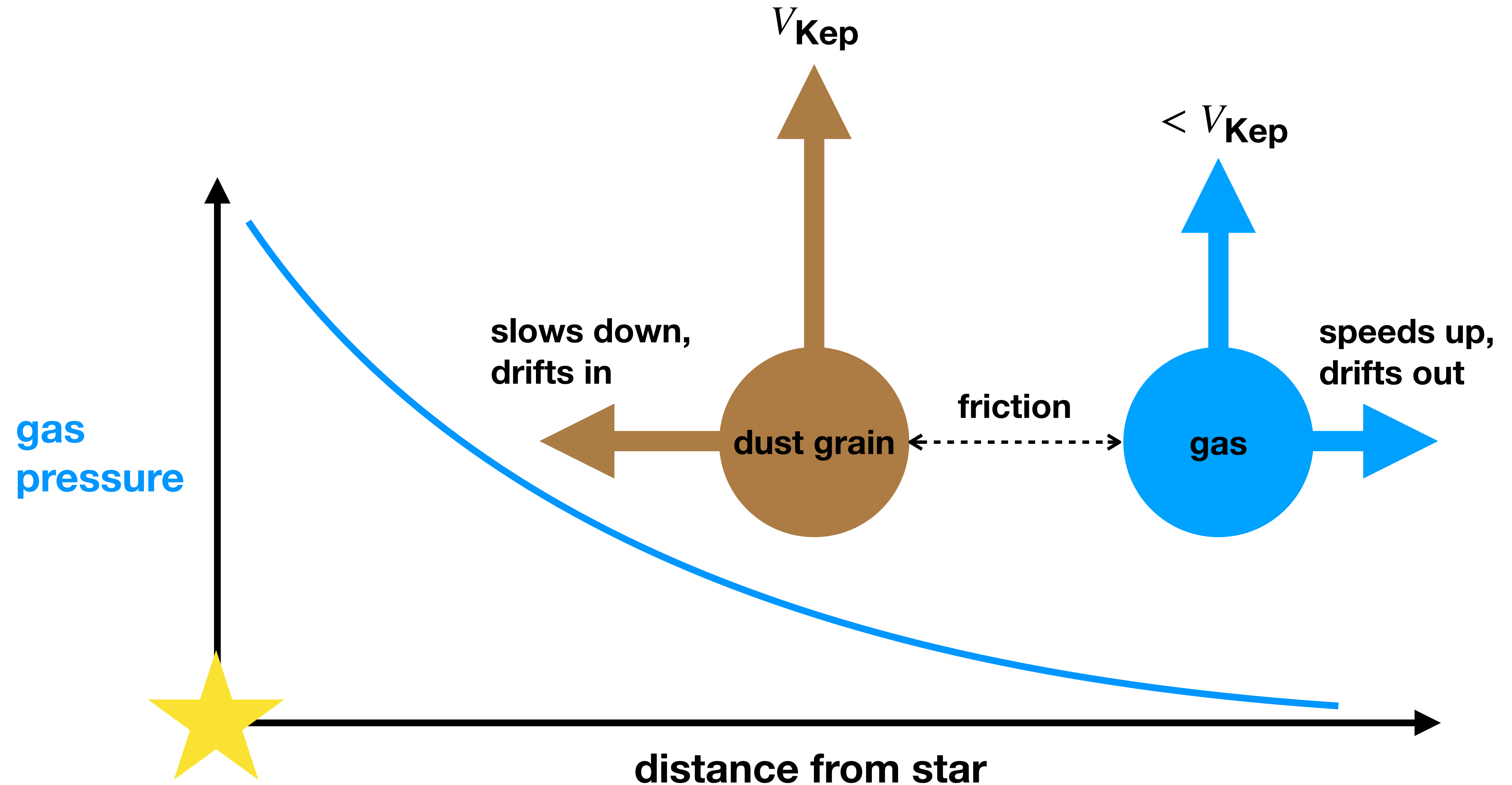




# Planetesimal formation via the SI



# Radial drift powers the classical SI



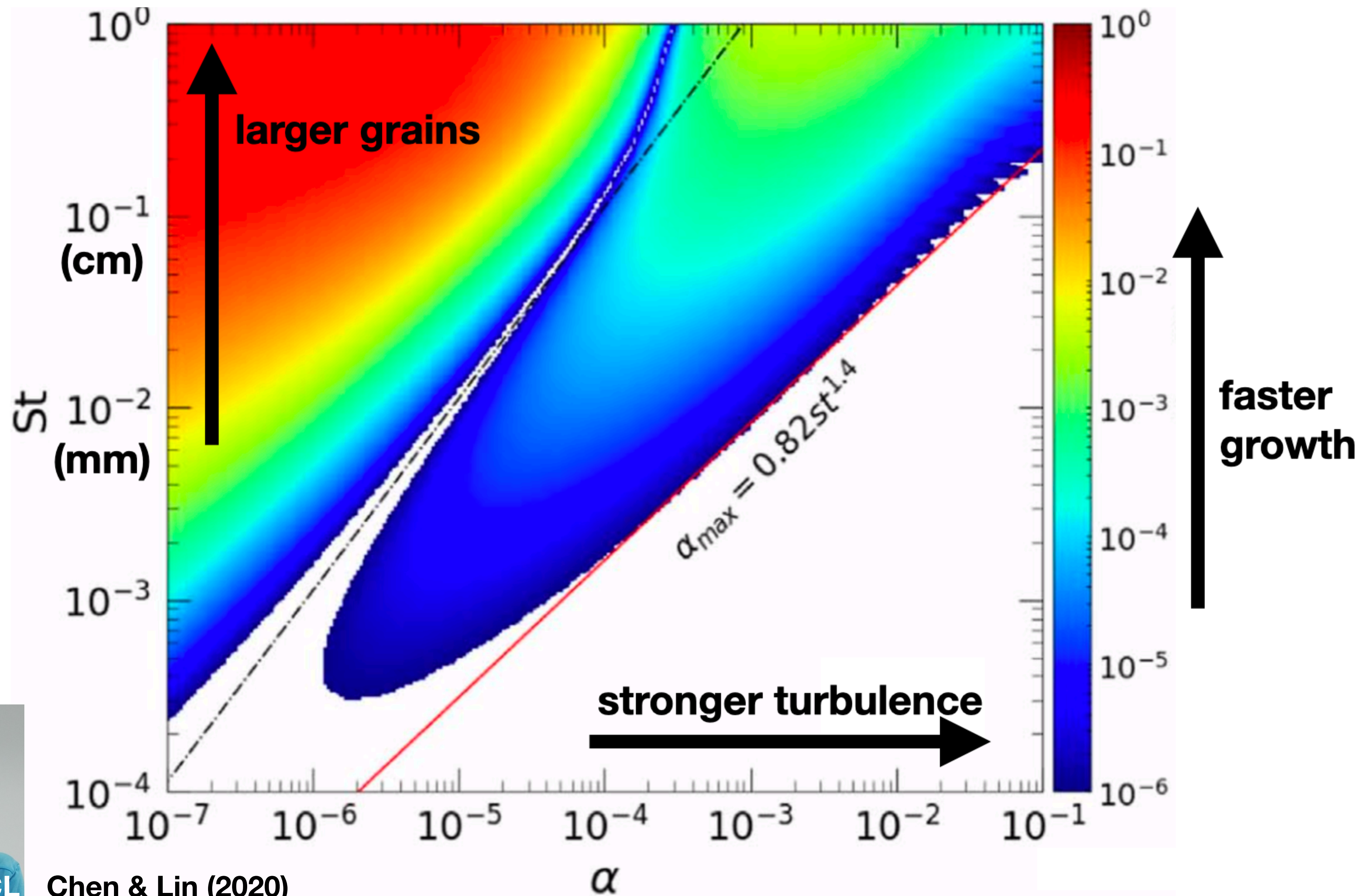


# The ideal SI

- **disk is non-turbulent** → **Chen & Lin (2020)**
- **disk has no vertical structure** → **Lin (2021)**
- **disk is unmagnetized** → **Lin & Hsu (2022)**  
**Hsu & Lin (in press)**



# SI is easily killed by viscosity

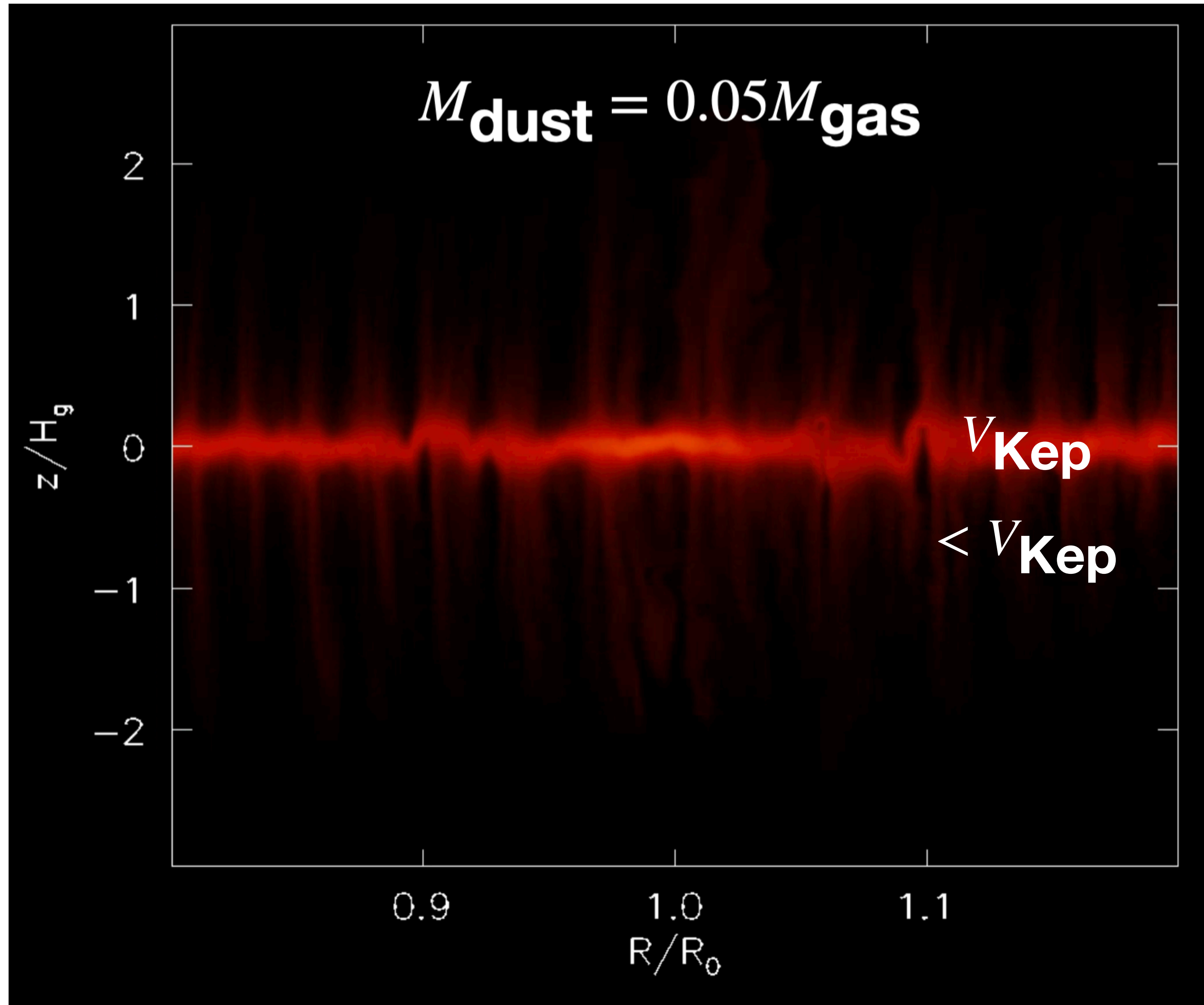


@UCL

Chen & Lin (2020)



# Stratified dust layers



$$\frac{\partial \Omega}{\partial z} \neq 0$$



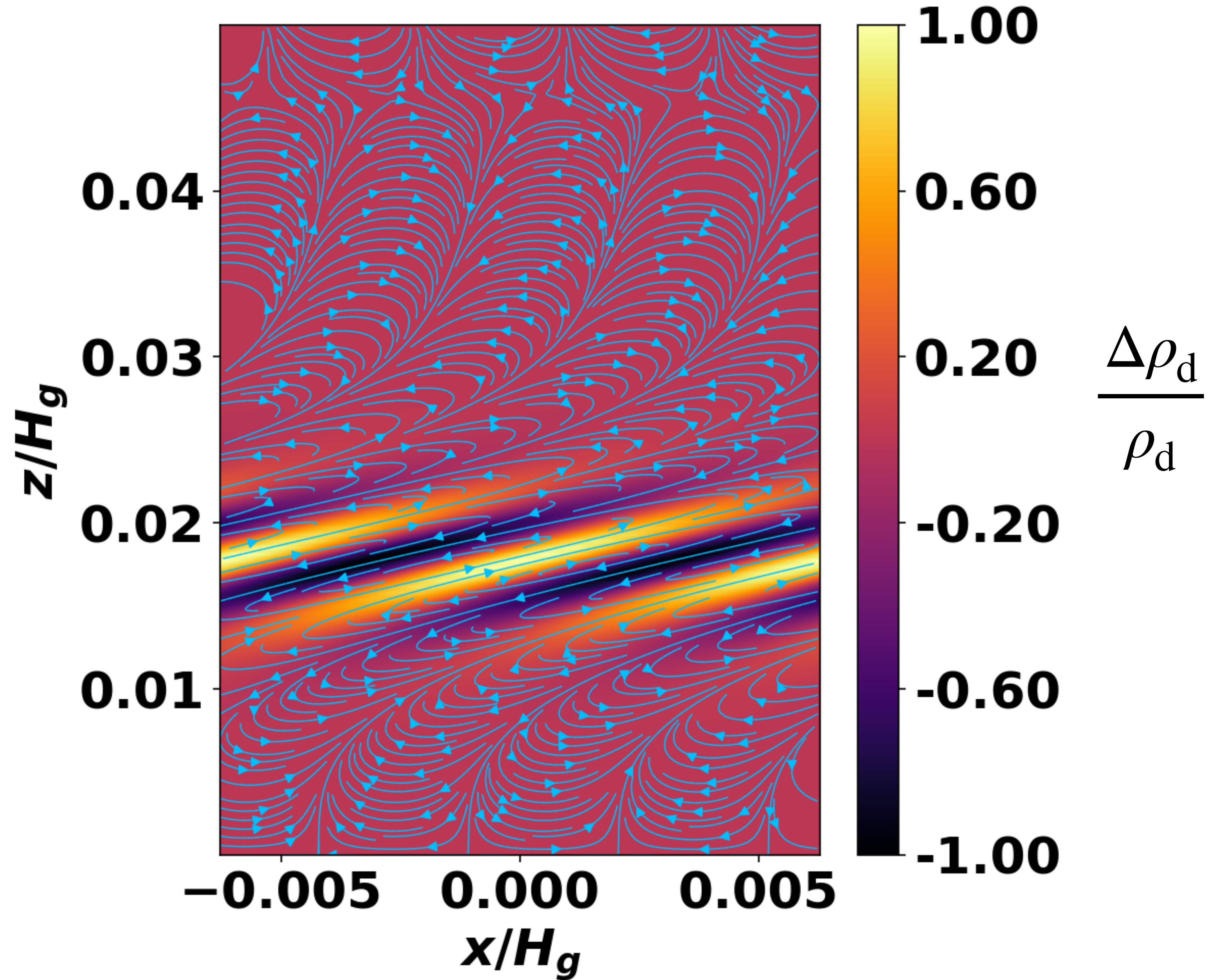
@ASIAA

Lehmann & Lin (2022): Dust settling in VSI-turbulent disks



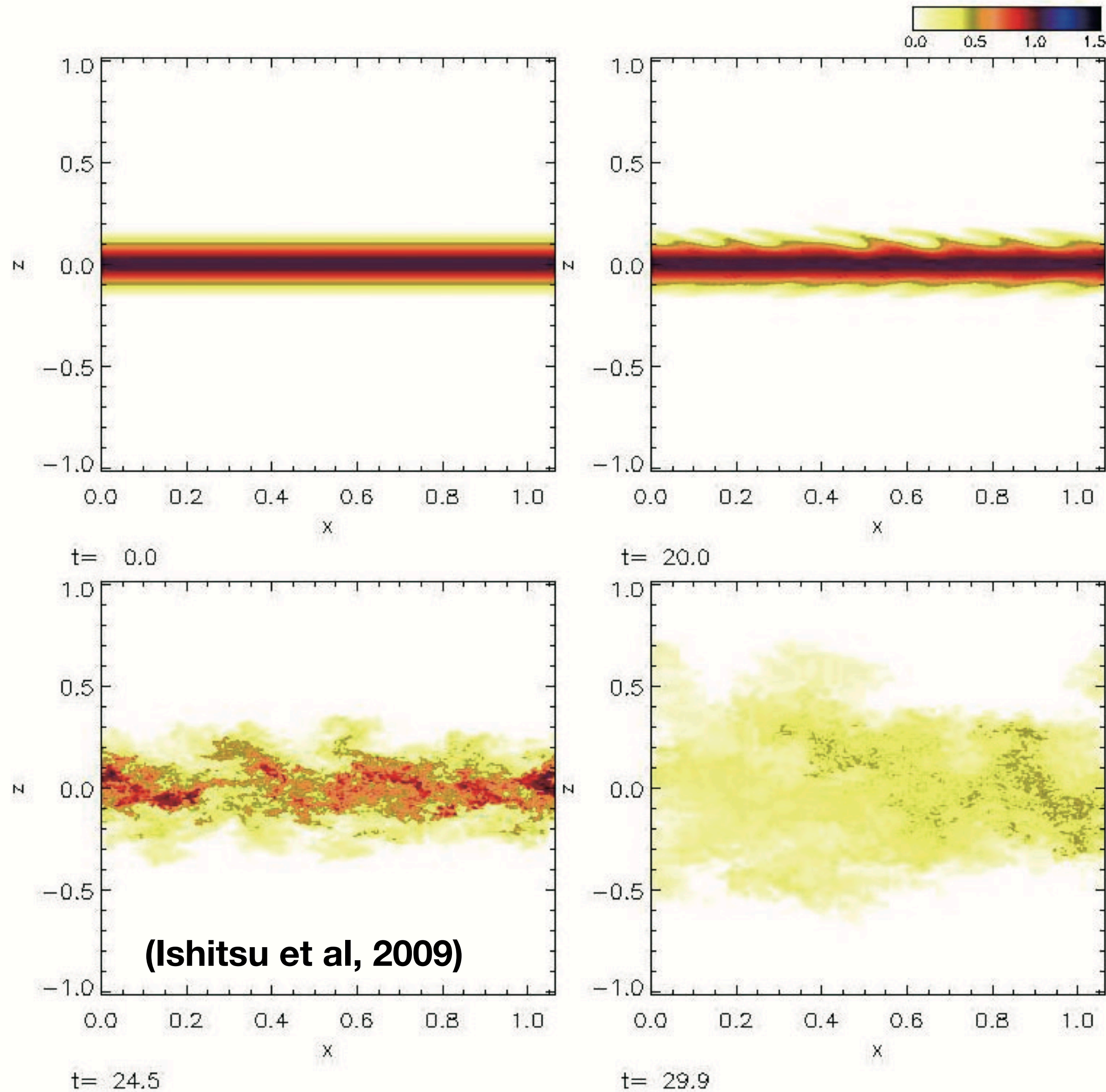
# “Vertically shearing SI” in stratified disks

$$S_{\text{grow}} \sim \Omega$$





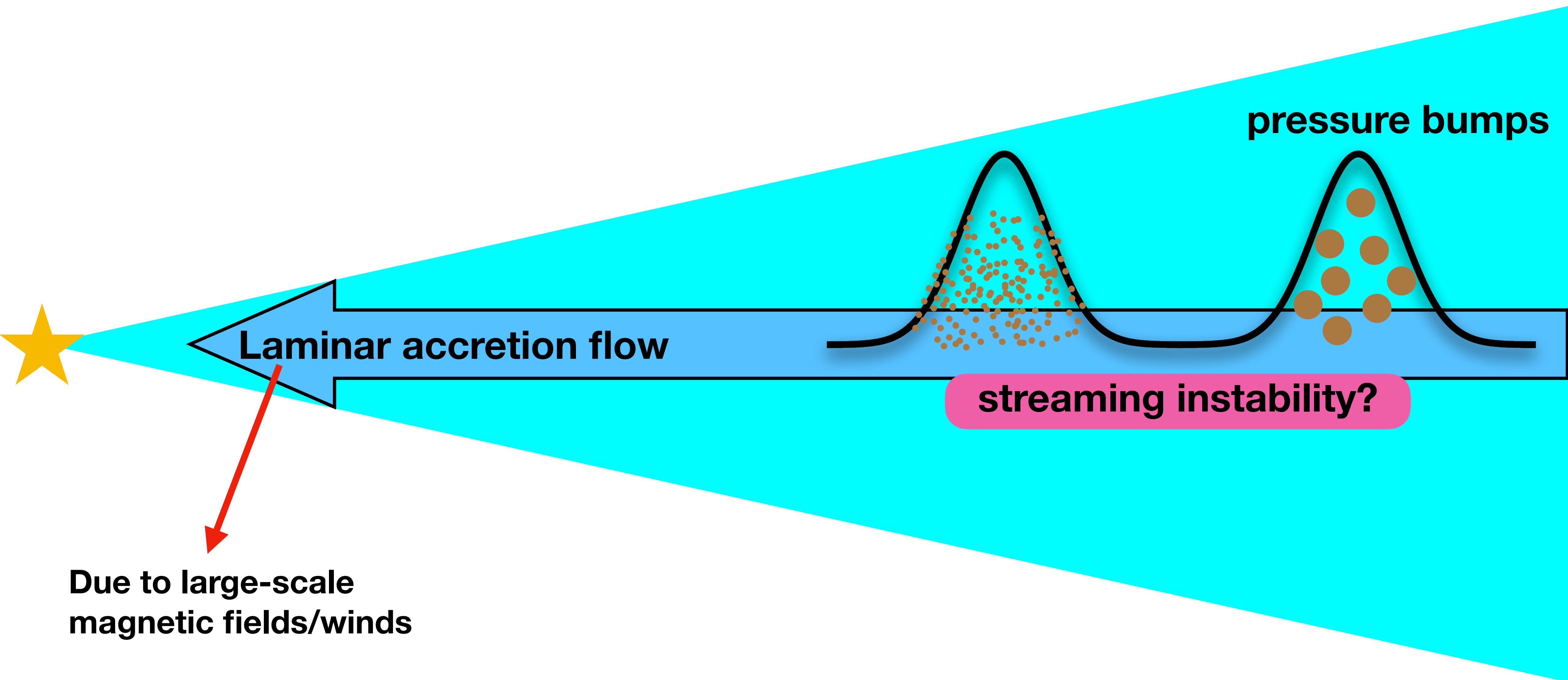
# Vertically shearing SIs grow fast but...



dust layer  
dispersed



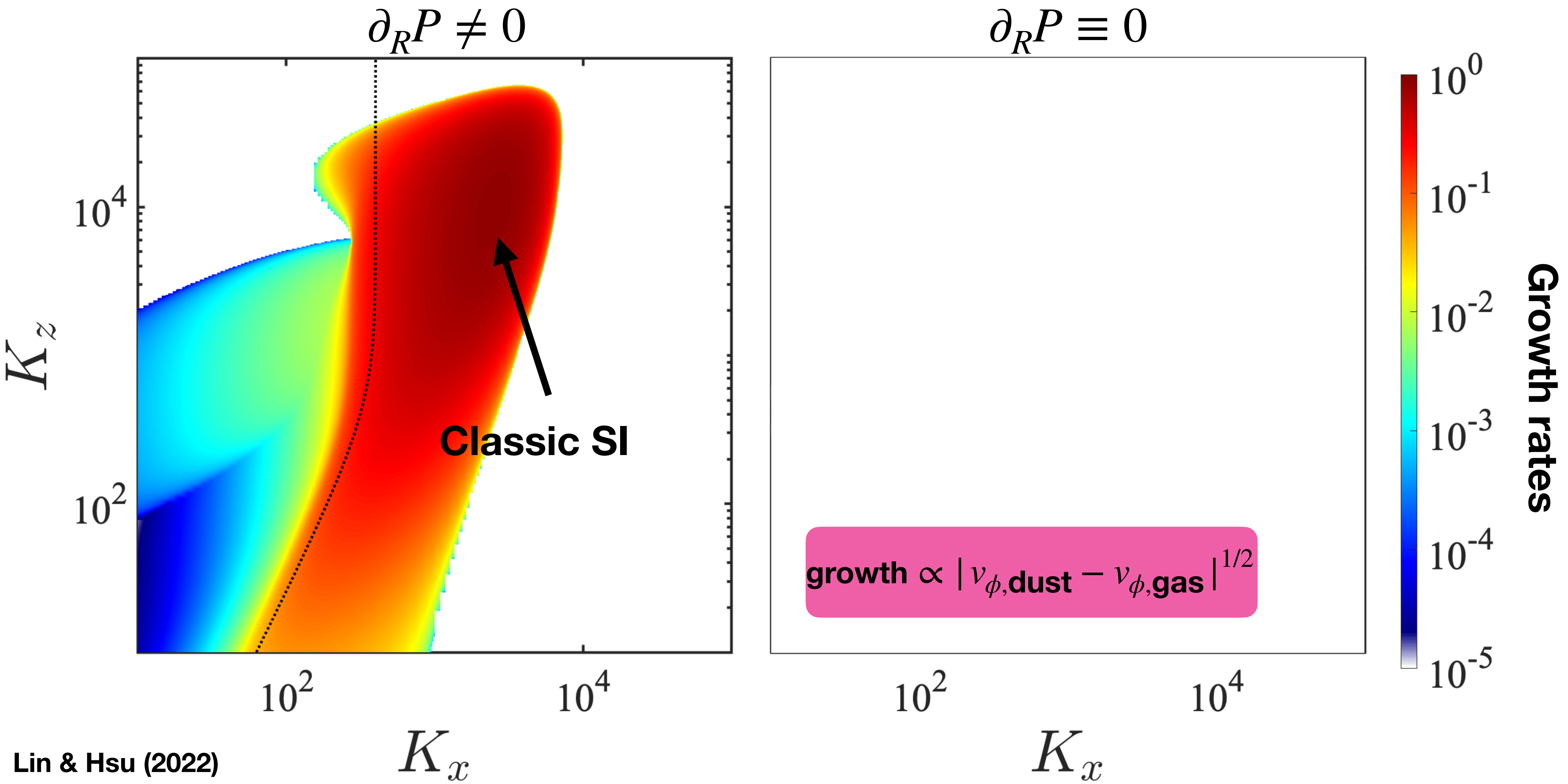
# Can modern disk models help?



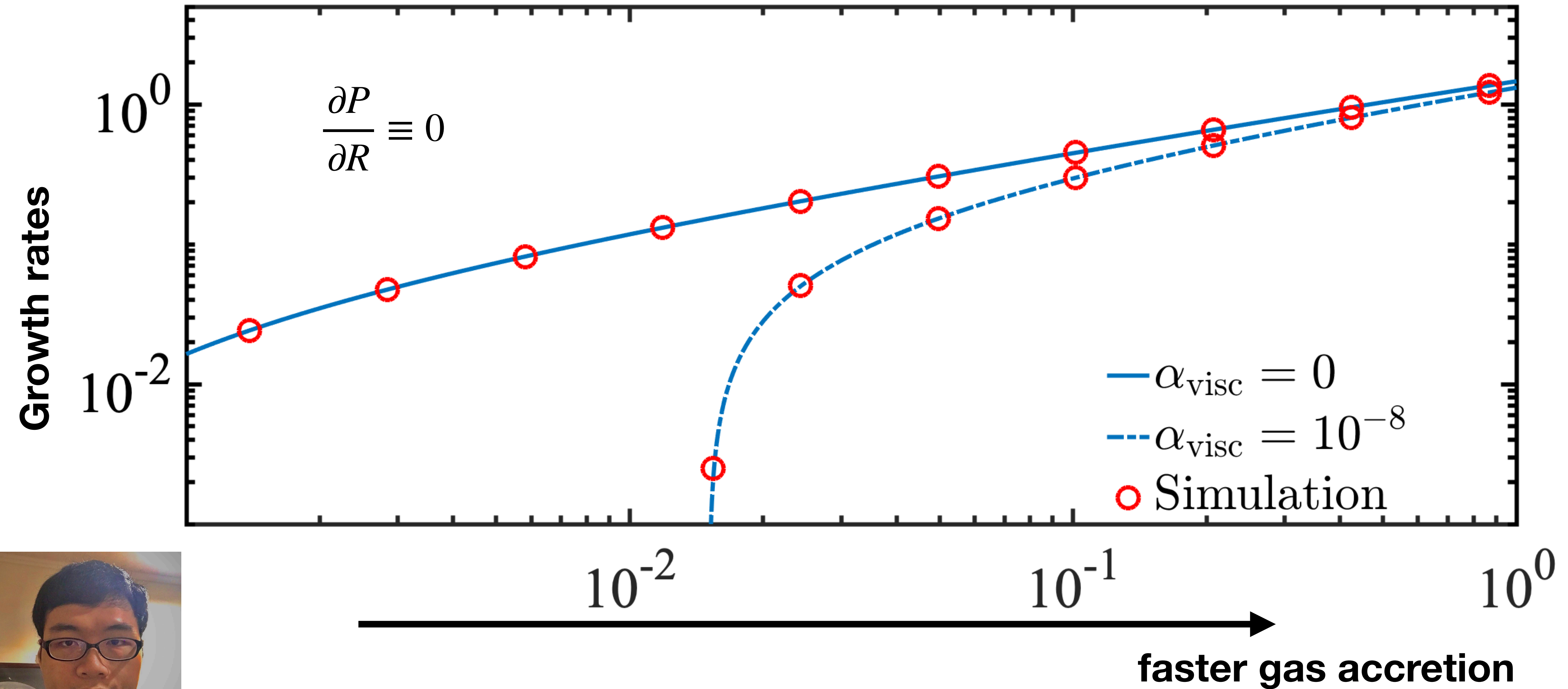
(e.g. Riols et al. 2020, Cui & Bai 2021)



# SI without pressure gradients



# Verification by FARGO3D



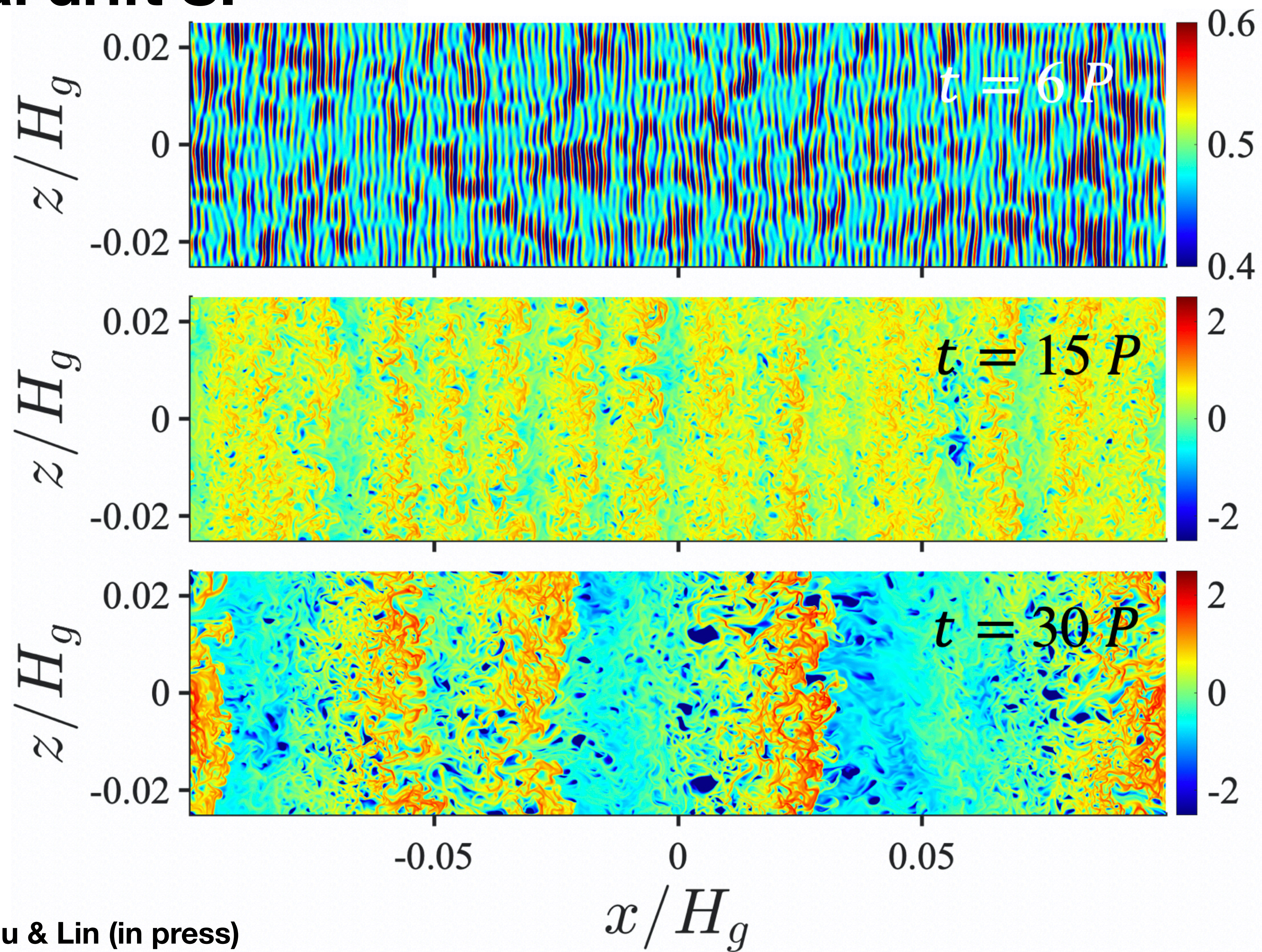
@UVa

Hsu & Lin (in press)



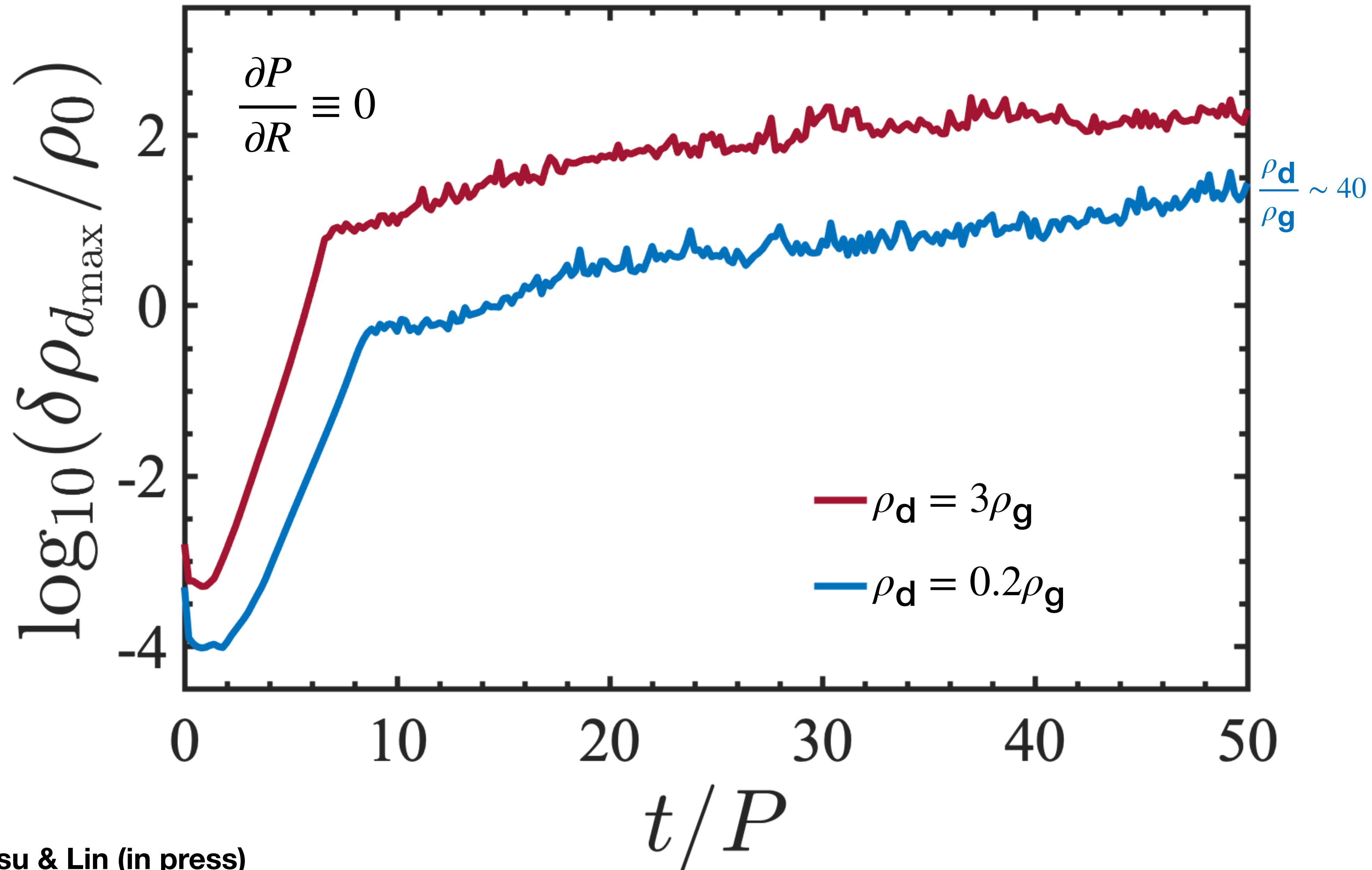
# Azimuthal drift SI

$$\frac{\partial P}{\partial R} \equiv 0$$



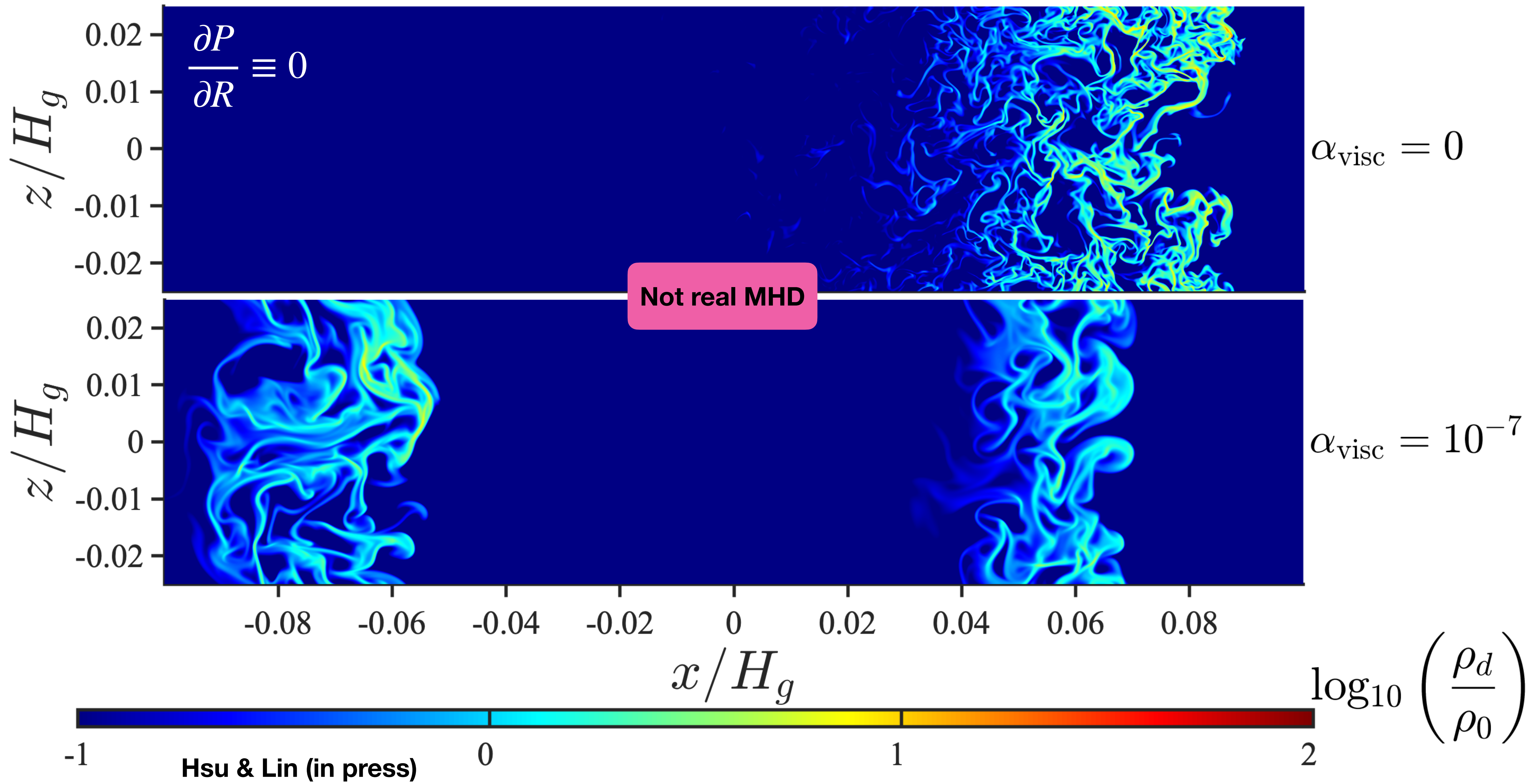


# Dust concentrates even when $\rho_d < \rho_g$ initially



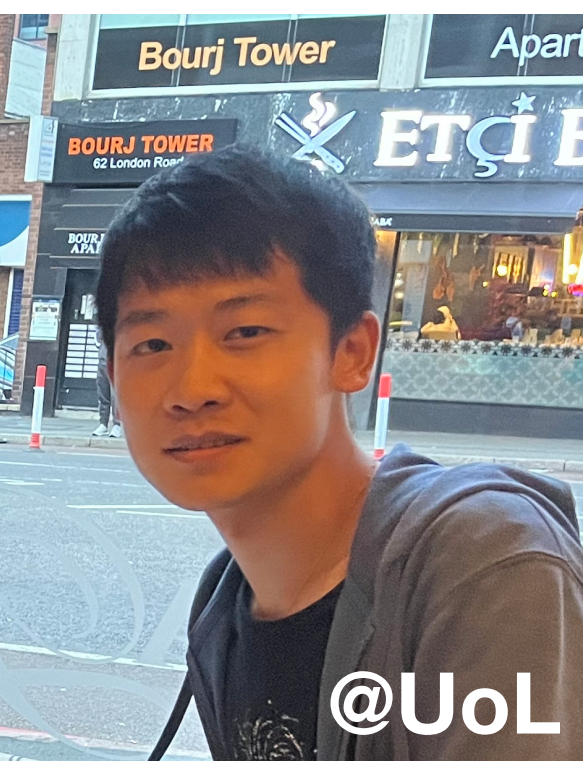
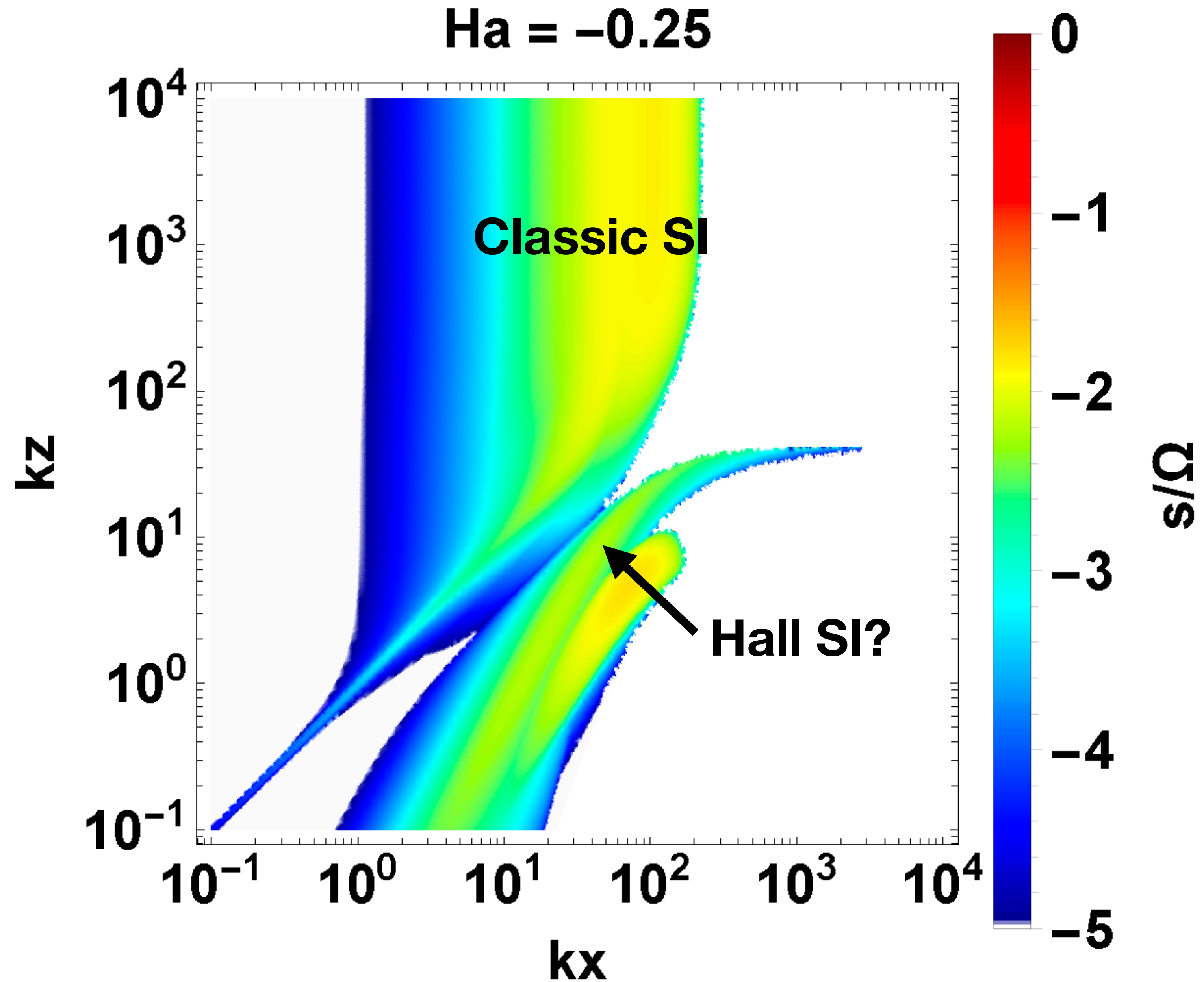


# Azimuthal drift SI, $\epsilon = 0.2, t = 50P$





# SI + live, non-ideal MHD



@UoL

Wu & Lin (in prep)



# Summary

- SI is easily killed by viscosity
- Vertically shearing SIs may disperse dust layers
- Azimuthal drift SI: works without pressure gradients, concentrates dust even at low metallicities

[Astrophysics](#) > [Earth and Planetary Astrophysics](#)

arXiv:2209.06784 (astro-ph)

[Submitted on 14 Sep 2022]

**Nonlinear evolution of streaming instabilities in accreting protoplanetary disks**

[Chun-Yen Hsu](#) (ASIAA), [Min-Kai Lin](#) (ASIAA, NCTS Physics Division)

On arXiv today!

**Thank you**  
 **@linminkai**